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Heat Transfer in the Flow of a Cold, Axisymmetric Jet Over a Hot Sphere

The heat-transfer characteristics of thin film flow over a hot sphere resulting from a cold vertical jet of liquid falling onto the surface have been investigated. The underlying physical features have been illustrated by numerical solutions of high accuracy based on the modified Keller box method. The solutions for film thickness distribution are good agreement with those obtained approximately by using the Pohlhausen integral momentum technique and observed experimentally by using water as working fluid, thus providing a basic confirmation of the validity of the results presented. [DOI: 10.1115/1.4007980]

Keywords: large Reynolds numbers, Keller box method

1 Introduction

Comprehensive literature reviews on theoretical and experimental studies of the flow and heat-transfer characteristics for jet impingement on surfaces were made by Thome [1] and Ribatski and Jacobi [2] to reveal fully the past developments. Although many different jet characteristics were discussed, a great majority of such studies were dealt with jet impingement on a flat surface. Jet impingement on curved surfaces is very common in engineering applications. Nevertheless, too few research works considered the question.

After having examined the jet impingement on a flat plate [3], it is natural to move on to a closer inspection of the jet impingement on a sphere. The flow of a thin liquid film under gravity over a sphere occurs frequently in a variety of industrial heat-transfer applications, such as, heat exchange in coating operations [4,5], high precision wetted spheres [6], multiple sphere absorbers [7], and hydrofluidization systems [8]. In order to understand the operations and, in particular, the efficiency of these processes, it is important to have a detailed study of such flows. The assessment of heat-transfer characteristics in such settings is based on Nusselt theory. Unfortunately the absence of inertia in the theory leads to the prediction of zero heat transfer at the upper generator of an inundated sphere. The work that follows in part addresses this inconsistency. Any underlying methodology of solution may indicate how best to incorporate inertia into a detailed assessment of the heat-transfer characteristics of jet impingement on a sphere. Analyses of such flows [9-12] fail to distinguish between the effects of a thin, high speed jet as compared to a thick, low speed jet when each give rise to a common flow rate. Mitrovic [13] indicates that such a distinguishing capability is called for the development of theoretical models. However previous studies of thin film flow over a sphere were confined purely to the hydrodynamic problem. Gyure and Krantz [14] used a perturbation analysis for low Reynolds numbers. Gribben [15] obtained an approximation by using the Pohlhausen integral momentum technique [16], which assumed an approximate velocity profile across the thickness of the film. Hunt [17] obtained a numerical solution by using the modified Keller box method, which accommodated the outer, free boundary. Heat-transfer characteristics of the flow have not been considered.

In this paper, the heat transfer in the flow of a cold, axisymmetric jet over a hot sphere is investigated. The accurate and

comprehensive numerical solutions for both velocity and temperature distributions are obtained by using the modified Keller box method. The main idea of the Keller box method [18] is to replace higher derivatives by first derivatives through the introduction of additional variables, and to discretize the resultant differential system by centered-difference derivatives with a second order truncation error at midpoints of net rectangles.

2 Modeling

The problem to be examined concerns the film cooling which occurs when a cold vertically draining column strikes a hot sphere. Although a column of fluid draining under gravity is accelerated and thin at impact [19,20], it is reasonable to model the associated volume flow as a jet of uniform velocity U_0 , uniform temperature T_0 , and radius H_0 as is illustrated in Fig. 1. The notation $Q = \pi H_0^2 U_0$ is introduced for the flow rate and a film Reynolds number may be defined as $\text{Re} = U_0 a/\nu$ based on the sphere radius a.

3 Governing Equations

The flow under investigation has been modelled as a steady, axisymmetric flow of incompressible fluid. For the large Reynolds number case in which the flow is confined to a thin film, quantities vary much less in the streamwise direction than in the wall normal direction. In the absence of viscous dissipation, the resultant governing equations expressing conservation of mass, momentum and energy become very similar to the boundary layer equations and are consequently

$$\frac{1}{\sin\theta} \frac{\partial}{\partial\theta} (V_{\theta} \sin\theta) + \frac{1}{r} \frac{\partial}{\partial r} (r^2 V_{\rm r}) = 0 \tag{1}$$

$$\frac{V_{\theta}}{r}\frac{\partial V_{\theta}}{\partial \theta} + v_r\frac{\partial V_{\theta}}{\partial r} + \frac{V_rV_{\theta}}{r} = g\sin\theta - \frac{1}{\rho r}\frac{dP}{d\theta} + \nu \left[\frac{1}{r^2}\frac{\partial}{\partial r}\left(r^2\frac{\partial V_{\theta}}{\partial r}\right) + \frac{1}{r^2\sin\theta}\frac{\partial}{\partial \theta}\left(\sin\theta\frac{\partial V_{\theta}}{\partial \theta}\right) + \frac{2}{r^2}\frac{\partial V_r}{\partial \theta} - \frac{V_{\theta}}{r^2\sin^2\theta}\right]$$
(2)

$$\frac{v_{\theta}}{r}\frac{\partial T}{\partial \theta} + v_{r}\frac{\partial T}{\partial r} = \kappa \left[\frac{1}{r^{2}}\frac{\partial}{\partial r}\left(r^{2}\frac{\partial T}{\partial r}\right) + \frac{1}{r^{2}\sin\theta}\frac{\partial}{\partial\theta}\left(\sin\theta\frac{\partial T}{\partial\theta}\right)\right]$$
(3)

where (V_{θ}, V_{r}) are velocity components associated with spherical coordinates (θ, r) measured by the angular displacement from the

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Fig. 1 The vertical jet and resultant film for the sphere

top of the sphere and the radial distance from the center of the sphere, respectively.

In the specified physical setting, the equations are to be solved subject to the following conditions:

(i) The no slip boundary condition at the wall requires that

$$V_{\theta} = V_{\rm r} = 0$$
 on $r = a$, $0 \le \theta \le \pi$ (4)

(ii) The temperature at the wall is assumed constant as T_w , say

i.e.,
$$T = T_w$$
 on $r = a$, $0 \le \theta \le \pi$ (5)

(iii) On the free surface of the film, prescribed by $r = a + H(\theta)$, the shearing stress may be assumed negligible and consequently

$$r\frac{\partial}{\partial r}\left(\frac{V_{\theta}}{r}\right) + \frac{1}{r}\frac{\partial V_{r}}{\partial \theta} = 0 \quad \text{at} \quad r = a + H(\theta), \quad 0 \le \theta \le \pi$$
(6)

(iv) Similarly, in a film cooling environment such as water surrounded by air, it may be assumed that there is negligible heat flux on the free surface and hence that

$$\frac{\partial T}{\partial r} = 0$$
 at $r = a + H(\theta), \quad 0 \le \theta \le \pi$ (7)

(v) Once an overall flow rate $Q = \pi H_0^2 U_0$ has been prescribed, a conservation of volume flow constraint at any given θ station leads to the condition

$$2\pi \sin\theta \int_{a}^{a+H(\theta)} rV_{\theta}(\theta, r)dr = \text{constant}$$
$$= \pi H_{0}^{2}U_{0} \quad \text{for} \quad 0 \le \theta \le \pi$$
(8)

Under the assumption that the film thickness remains thin relative to a characteristic horizontal dimension, a boundary layer treatment of the equations leads to significant simplification.

The following nondimensional variables are introduced:

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$$x = \theta, \quad \bar{Y} = \frac{\mathrm{Re}^{1/2}(r-a)}{a}, \quad \bar{H}(x) = \frac{\mathrm{Re}^{1/2}H(\theta)}{a}$$
$$\bar{U} = \frac{V_{\theta}}{U_0}, \quad \bar{V} = \frac{\mathrm{Re}^{1/2}V_{\mathrm{r}}}{U_0}, \quad \phi = \frac{T-T_{\mathrm{w}}}{T_0 - T_{\mathrm{w}}}, \quad p = \frac{P}{\rho U_0^2}$$
(9)

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The big O notation, O, is used to describe the limiting behavior of a function for a very small argument. In the limit $\text{Re} \to +\infty$ with x remaining O(1) and after neglecting terms of $O((1/\text{Re}^{1/2}) + (H_0^2/a^2))$ compared with unity, the following equations are obtained:

$$\frac{\partial}{\partial x}(\bar{U}\sin x) + \frac{\partial}{\partial\bar{Y}}(\bar{V}\sin x) = 0$$
(10)

$$\bar{U}\frac{\partial\bar{U}}{\partial x} + \bar{V}\frac{\partial\bar{U}}{\partial\bar{Y}} = \frac{1}{\mathrm{Fr}}\sin x - \frac{dp}{dx} + \frac{\partial^2\bar{U}}{\partial\bar{Y}^2}$$
(11)

$$\Pr\left(\bar{U}\frac{\partial\phi}{\partial x} + \bar{V}\frac{\partial\phi}{\partial\bar{Y}}\right) = \frac{\partial^2\phi}{\partial\bar{Y}^2}$$
(12)

where $Pr = \nu/\kappa$ is the Prandtl number and $Fr = U_0^2/ag$ is the Froude number based on the jet velocity on its surface. In common with standard boundary layer theory, dp/dx implies that the pressure across the film remains constant. In the absence of external pressure gradients and with zero shear assumed on the free surface, the pressure term in Eq. (11) is identically zero.

The boundary conditions now read

(

i)
$$\bar{U} = \bar{V} = 0$$
 on $\bar{Y} = 0, \quad 0 \le x \le \pi$ (13)

ii)
$$\phi = 0$$
 on $\bar{Y} = 0, \quad 0 \le x \le \pi$ (14)

(iii)
$$\frac{\partial U}{\partial \bar{Y}} = 0$$
 at $\bar{Y} = \bar{H}(x), \quad 0 \le x \le \pi$ (15)

iv)
$$\frac{\partial \phi}{\partial \bar{Y}} = 0$$
 at $\bar{Y} = \bar{H}(x), \quad 0 \le x \le \pi$ (16)

v)
$$\int_{0}^{\bar{H}(x)} \bar{U}d\bar{Y} = \frac{R_{\rm e}^{1/2}H_0^2}{2a^2\sin x} \quad \text{for} \quad 0 \le x \le \pi$$
(17)

4 Numerical Solutions

The continuity Eq. (10) can be eliminated by introducing a stream function ψ defined by

$$\bar{U} = \frac{1}{\sin x} \frac{\partial \psi}{\partial \bar{Y}}, \quad \bar{V} = -\frac{1}{\sin x} \frac{\partial \psi}{\partial x}$$
 (18)

Owing to the geometry, $\overline{H}(x)$ is singular at x = 0 and $x = \pi$. To remove this singularity, y and h(x) are introduced and given by

$$y = \sin xY, \quad h(x) = \sin xH(x) \tag{19}$$

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Substituting Eqs. (18) and (19) into Eqs. (10)–(17) gives

$$\frac{\partial^3 \psi}{\partial y^3} + \frac{1}{\operatorname{Fr}\sin x} = \left(\frac{1}{\sin^2 x}\right) \left(\frac{\partial \psi}{\partial y} \frac{\partial^2 \psi}{\partial x \partial y} - \frac{\partial \psi}{\partial x} \frac{\partial^2 \psi}{\partial y^2}\right)$$
(20)

$$\frac{\partial^2 \phi}{\partial y^2} = \left(\frac{P_{\rm r}}{\sin^2 x}\right) \left(\frac{\partial \psi}{\partial y} \frac{\partial \phi}{\partial x} - \frac{\partial \psi}{\partial x} \frac{\partial \phi}{\partial y}\right) \tag{21}$$

subject to boundary conditions

$$\psi = 0, \quad \frac{\partial \psi}{\partial y} = 0, \quad \phi = 0 \quad \text{on} \quad y = 0, \quad 0 \le x \le \pi$$
 (22)

$$\psi = \frac{\operatorname{Re}^{1/2} H_0^2}{2a^2}, \quad \frac{\partial^2 \psi}{\partial y^2} = 0, \quad \frac{\partial \phi}{\partial y} = 0 \quad \text{on} \quad y = h(x), \quad 0 \le x \le \pi$$
(23)

$$h = \frac{\operatorname{Re}^{1/2} H_0^2}{2a^2}, \quad \psi = y, \quad \phi = 1 \quad \text{on} \quad x = 0, \quad 0 < y \le \frac{\operatorname{Re}^{1/2} H_0^2}{2a^2} \quad (24)$$

where the initial condition (24) appears due to the original initial condition

$$H = -a + \sqrt{a^2 + \frac{H_0^2}{\sin \theta}}, \quad V_\theta = U_0, \quad T = T_0 \quad \text{on} \quad \theta = 0,$$
$$a < r \le \sqrt{a^2 + \frac{H_0^2}{\sin \theta}} \tag{25}$$

The main idea of the Keller box method [18] is to replace higher derivatives by first derivatives through the introduction of additional variables, and to discretize the resultant differential system by centered-difference derivatives with a second order truncation error at midpoints of net rectangles. Some progress [21–25] has been made in investigating various problems of a liquid jet impinging on a solid surface. In anticipation of the use of a Keller box method and its attractive extrapolation features the differential system (20)–(24) is recast as the following first order system:

$$\frac{\partial \psi}{\partial y} = u$$
$$\frac{\partial u}{\partial y} = \bar{v}$$
$$\frac{\partial \bar{v}}{\partial y} = -\frac{1}{F_{\rm r} \sin x} + \left(\frac{1}{\sin^2 x}\right) \left(u \frac{\partial u}{\partial x} - \bar{v} \frac{\partial \psi}{\partial x}\right) \qquad (26)$$
$$\frac{\partial \phi}{\partial y} = \bar{w}$$
$$\frac{\partial \bar{w}}{\partial y} = \left(\frac{\Pr}{\sin^2 x}\right) \left(u \frac{\partial \phi}{\partial x} - \bar{w} \frac{\partial \psi}{\partial x}\right)$$

whose boundary conditions are

$$\psi = 0, \quad u = 0, \quad \phi = 0 \quad \text{on} \quad y = 0, \quad 0 \le x \le \pi$$

$$\psi = \frac{\operatorname{Re}^{1/2} H_0^2}{2a^2}, \quad \bar{v} = 0, \quad \bar{w} = 0 \quad \text{on} \quad y = h(x), \quad 0 \le x \le \pi$$

$$(27)$$

$$h = \frac{\operatorname{Re}^{1/2} H_0^2}{2a^2}, \quad \psi = y, \quad \phi = 1 \quad \text{on} \quad x = 0, \quad 0 < y \le \frac{\operatorname{Re}^{1/2} H_0^2}{2a^2}$$

If these equations were used as the basis of solution there would be strong comparisons between the associated algorithm and that developed in Ref. [26]. However the expectation of an initial Blasius boundary layer within the film can be assimilated

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into the solution scheme by further transformations. The underlying methodology of solution nevertheless remains the same. In each case the discretisation is aimed at generating a simultaneous system of nonlinear equations which can be solved by Newton iteration.

According to the nondimensional transformation, the boundary layer thickness grows like $x^{3/2}$ for small *x* in the *y* direction.

The following coordinate transformation, what simultaneously maps the film thickness onto the unit interval and removes the Blasius singularity at the origin, is introduced:

$$x = \xi^{2/3}, \quad y = \frac{\xi \eta h}{\xi + 1 - \eta}$$
 (28)

The dependent variables are transformed as

$$\psi = \frac{\xi}{\xi + 1 - \eta} f, \quad \bar{v} = \frac{\xi + 1 - \eta}{\xi} v, \quad \bar{w} = \frac{\xi + 1 - \eta}{\xi} w \quad (29)$$

The equations to be solved now read

$$f_{\eta} = \frac{(1+\xi)hu}{\xi+1-\eta} - \frac{f}{\xi+1-\eta}$$
$$u_{\eta} = \frac{(1+\xi)hv}{\xi+1-\eta}$$
$$v_{\eta} = \frac{v}{\xi+1-\eta} - \frac{\xi^{2}(1+\xi)h}{\operatorname{Fr}(\xi+1-\eta)^{3}\operatorname{sin}(\xi^{2/3})}$$
$$- \frac{3(1-\eta)\xi^{4/3}(1+\xi)hfv}{2(\xi+1-\eta)^{4}\operatorname{sin}^{2}(\xi^{2/3})} + \frac{3\xi^{7/3}(1+\xi)h}{2(\xi+1-\eta)^{3}\operatorname{sin}^{2}(\xi^{2/3})}(uu_{\xi} - vf_{\xi})$$
(30)

$$\phi_{\eta} = \frac{(\xi - 1)^{4/3}}{\xi + 1 - \eta}$$
$$w_{\eta} = \frac{w}{\xi + 1 - \eta} - \frac{3 \operatorname{Pr}(1 - \eta) \xi^{4/3} (1 + \xi) h f w}{2(\xi + 1 - \eta)^{4} \operatorname{sin}^{2}(\xi^{2/3})} + \frac{3 \operatorname{Pr} \xi^{7/3} (1 + \xi) h}{2(\xi + 1 - \eta)^{3} \operatorname{sin}^{2}(\xi^{2/3})} (u \phi_{\xi} - w f_{\xi})$$

subject to

$$f = 0, \quad u = 0, \quad \phi = 0 \quad \text{on} \quad \eta = 0, \quad 0 \le \xi \le \pi^{3/2}$$

$$f = \frac{\operatorname{Re}^{1/2}H_0^2}{2a^2}, \quad v = 0, \quad w = 0 \quad \text{on} \quad \eta = 1, \quad 0 \le \xi \le \pi^{3/2} \quad (31)$$

$$h = \frac{R_e^{1/2}H_0^2}{2a^2}, \quad f = f_0(\eta), \quad \phi = \phi_0(\eta) \quad \text{on} \quad \xi = 0, \quad 0 < \eta \le 1$$

where the initial profiles $f_0(\eta)$ and $\phi_0(\eta)$ are found by putting $\xi = 0$ and $h = (\text{Re}^{1/2}H_0^2)/(2a^2)$ into Eq. (30) and solving subject to conditions $f = u = \phi = 0$ at $\eta = 0$ and u = 1, $\phi = 1$ at $\eta = 1$.

The parabolic system of equations and boundary conditions (30) and (31) has been solved by marching in the ξ -direction based on a modification of the Keller box method. A nonuniform grid is placed on the domain $\xi \ge 0$, $0 \le \eta \le 1$ and the resultant difference equations are solved by Newton iteration. Solutions are obtained on different sized grids and Richardson's extrapolation used to produce results of high accuracy.

It is worth to mention to this end that the general equations describing parabolic free boundary problems arising from thin film flows [26] are

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Table 1 Film thickness, free surface velocity and temperature for the sphere with Fr = 1, γ = 0.5, and Pr = 2

x	Film thickness $\bar{h}(x)$	Free surface velocity $\bar{u}(x, \bar{h}(x))$	Free surface temperature $\bar{\phi}(x, \bar{h}(x))$
0.000	0.500	1.000	1.000
0.218	0.585	1.023	1.000
0.396	0.678	1.063	1.000
0.587	0.775	0.997	0.964
0.786	0.891	0.864	0.831
0.994	1.007	0.756	0.654
1.208	1.086	0.695	0.484
1.427	1.128	0.667	0.343
1.595	1.141	0.658	0.260
1.823	1.138	0.659	0.180
1.996	1.122	0.667	0.139
2.290	1.072	0.695	9.760×10^{-2}
2.590	0.994	0.744	7.779×10^{-2}
3.018	0.861	0.823	7.022×10^{-2}
π	0.836	0.831	7.013×10^{-2}

$$\frac{\partial^3 \psi}{\partial y^3} + F(x) = G^2(x) \left(\frac{\partial \psi}{\partial y} \frac{\partial^2 \psi}{\partial x \partial y} - \frac{\partial \psi}{\partial x} \frac{\partial^2 \psi}{\partial y^2} \right)$$
(32)

$$\frac{\partial^2 \bar{\phi}}{\partial y^2} = P_{\rm r} G^2(x) \left(\frac{\partial \psi}{\partial y} \frac{\partial \bar{\phi}}{\partial x} - \frac{\partial \psi}{\partial x} \frac{\partial \bar{\phi}}{\partial y} \right) \tag{33}$$

subject to boundary conditions

$$\psi = 0, \quad \frac{\partial \psi}{\partial y} = 0, \quad \bar{\phi} = 0 \quad \text{on} \quad y = 0, \quad 0 \le x \le x_{s}$$
 (34)

$$\psi = \gamma$$
, $\frac{\partial^2 \psi}{\partial y^2} = 0$, $\frac{\partial \bar{\phi}}{\partial y} = 0$ at $y = \bar{h}(x)$, $0 \le x \le x_s$ (35)

$$\bar{h} = \gamma, \quad \psi = y, \quad \bar{\phi} = 1 \quad \text{on} \quad x = 0, \quad 0 < y \le \gamma$$
 (36)

Accordingly, the variables are transformed from (x, y) to (ξ, η) by using

$$x = \xi^{\alpha}, \quad y = \frac{\xi \eta h}{\xi + 1 - \eta}$$
(37)

The dependent variables are transformed as

$$\psi = \frac{\xi}{\xi + 1 - \eta} f, \quad \bar{u} = \frac{1}{(1 + \xi)^{\beta}} u, \quad \bar{v} = \frac{\xi + 1 - \eta}{\xi (1 + \xi)^{2\beta}} v,$$
$$\bar{\phi} = \phi, \quad \bar{w} = \frac{\xi + 1 - \eta}{\xi (1 + \xi)^{\beta}} w, \quad \bar{h} = (1 + \xi)^{\beta} h \tag{38}$$

Here, for the sphere in this paper, the relevant physical parameters in the parabolic system of general equations, boundary conditions and coordinate transformation (32)–(38) describing free boundary problems arising from thin film flows should be chosen as the gravity-related parameter $F(x) = 1/\text{Fr} \sin x$, the geometryrelated parameter $G(x) = 1/\sin x$, the domain-related parameter $x_s = \pi$, the curvature-related parameter $\gamma = R_e^{1/2}H_0^2/2a^2$, the thickness-related parameter $\alpha = \frac{2}{3}$, and the similarity-related parameter $\beta = 0$. A full account of the numerical method and the details of implementation are beyond the scope of this paper and is reported separately [26]. The detailed numerical method procedure for this case is fully discussed in Ref. [27]. The solution scheme was successfully tested against previously reported results [28–33].

5 Results

A typical run has a coarse grid of dimensions 60×48 in the (ξ, η) domain with each cell being divided into 1, 2, 3, and 4 subcells respectively. Because of the coordinate singularity at $\xi = 0$, $\eta = 1$, a nonuniform grid is employed and given by $\xi = \overline{\xi}^{1.75}$, $\eta = 1 - (1 - \overline{\eta})^{1.5}$ where $\overline{\xi}$ and $\overline{\eta}$ are uniform. When $\Delta \overline{\xi} \equiv \frac{1}{59} \pi^{6/7}$ and $\Delta \overline{\eta} \equiv \frac{1}{47}$, this gives $\Delta \xi \sim 0.004$ and $\Delta \eta \sim 0.003$ near the singularity, which is sufficiently small to give good accuracy. From the convergence of the extrapolation process, the absolute error is 6×10^{-7} . A typical set of numerical data is presented in Table 1.

In Fig. 2, the numerical solution for the film thickness distribution over the sphere is compared with Gribben's approximation [15] for $F_r = 1$ and $\gamma = 0.5$. The agreement is seen to be surprisingly good. Figures 3–9 depict the flavour of the numerical results. Figures 3 and 4, Figs. 5 and 6, and Figs. 7–9 show film thickness, free surface velocity, and free surface temperature, respectively, for various cases.

It is worth to mention that the author's previous published results on a flat plate [3] can be regarded as a limiting case of the present new ones on a sphere, when the focus is only on the near field around jet impingement point $x \to 0$ with negligible gravity effect $Fr \to \infty$ and negligible curvature effect $\gamma \to 0.5$. This can be evidenced by making the comparison between Figs. 3, 5, and 8 for the upper hemispherical surface $0 \le x \le \frac{\pi}{2}$ at $Fr \ge 5$ and the counterpart figures in Ref. [3] (collectively as shown in Fig. 10 for the sake of easy reference), but for the lower hemispherical



Fig. 2 Film thickness for the numerical solution and Gribben's approximation at Fr = 1 and $\gamma = 0.5$

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Fig. 3 Film thickness for various Froude numbers at $\gamma = 0.5$



Fig. 4 Film thickness for various values of the parameter γ at Fr = 1



Fig. 5 Free surface velocity for various Froude numbers at $\gamma=0.5$

surface $\frac{\pi}{2} \le x \le \pi$ at Fr < 5, the characteristics are totally different between two cases.

For the sphere case, the velocity of the flow is controlled by two opposing forces, viscosity trying to slow it down (less heat transfer) and gravity trying to speed it up (more heat transfer). The viscous component of force affecting the flow is greatest near

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Fig. 6 Free surface velocity for various values of the parameter γ at Fr = 1



Fig. 7 Free surface temperature for various Prandtl numbers at Fr = 1 and γ = 0.5



Fig. 8 Free surface temperature for various Froude numbers at $\gamma = 0.5$ and Pr = 2

 $x = \frac{\pi}{2}$ and least near x = 0 and $x = \pi$. Figure 5 show that at Fr = 1, free surface velocity initially has a slight increase, followed by a sharp decrease as viscosity starts to dominate, and finally a gradual increase is observed as the bottom of the sphere is approached. This corresponds to the up-down situation for the film thickness in Figs. 2–4, and the decline at the different rates for the free surface temperature in Figs. 7–9. This phenomenon was observed experimentally [34] by using water as working fluid for jet impingement Reynolds number ranging from the order of 10⁴ to 10⁶. As Fr decreases, the effect of gravity increases and hence the thin film thickness, high velocity and low temperature

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Fig. 9 Free surface temperature for various values of the parameter γ at Fr = 1 and Pr = 2



Fig. 10 Film thickness, free surface temperature and free surface velocity for the flat plate case [3]



Fig. 11 Nusselt number for the sphere with Fr = 1, γ = 0.5, and Pr = 2

appear corresponding to the small Fr values. As γ decreases, the amount of fluid in the impinging jet decreases and the ensuing film becomes thinner. The effect of viscosity increases and hence the low velocity and low temperature appear corresponding to the small γ values. Where γ is greater than a certain value, e.g., $\gamma = 2$ for Fr = 1, Fig. 6 shows no decrease for the velocity over the sphere. As Pr increases, the temperature decrease becomes more gradual.

It is worth to mention to this end that the most significant film cooling design factor is the heat transfer across the film. This can be presented most conveniently by introduction of the Nusselt number defined as

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$$\operatorname{Nu} = \frac{\partial \phi}{\partial y}\Big|_{y=0} = \frac{x^{3/2} + 1}{x^{3/2}} w\Big|_{\eta=0}$$
(39)

A typical result of the Nusselt number, Nu, for the sphere with $F_r = 1$, $\gamma = 0.5$, and $P_r = 2$ is illustrated in Fig. 11. It is shown that the Nusselt number, Nu, increases monotonically as the rate of $x^{3/2}$ as approaching to the top of the sphere. This is in contrast to models based solely on a balance of viscous and gravitational terms, which necessarily predict zero Nusselt number at the top of the sphere. In contrast to the classical Nusselt theory, the film inertia may generate significant heat transfer (high Nusselt number) at the top of the sphere.

6 Concluding Remarks

A detailed examination of a jet over a sphere has been performed. The accurate and comprehensive numerical solutions for establishing the flow and heat-transfer characteristics of a cold axisymmetric jet over a hot sphere have been presented by modifying the Keller box method to accommodate the outer, free boundary. The gross features of such flows have been illustrated over a range of representative parameter values. These indicate the underlying features of the developing film thickness, velocity and temperature distributions. The comparison with the experimental observation [34] by using water as working fluid for jet impingement Reynolds number ranging from the order of 10⁴ to 10⁶ is to indicate that the current numerical solution may be carried over with confidence to the sphere inundation problem, thus providing a basis of comparison with Mitrovic's experimental results [13]. The work also provides the basis for reassessing condensation drainage and inundation flows; recognizing that in contrast to Nusselt theory, the inertia of inundating film may generate significant heat transfer at the top of a flooded sphere. In particular such an implementation of the present method is likely to yield significant heat transfer at the top of an inundated sphere. This is in contrast to models based solely on a balance of viscous and gravitational terms, which necessarily predict zero heat transfer at the upper generator. In a practical setting, appropriate parameter values may be evaluated and the design characteristics readily identified from the numerical solutions. In practice, it is not obvious that uniform wetting of the sphere would occur. Instabilities may distort or even disrupt such a uniform distribution, that is, U_0 may be not considered as average velocity rather than uniform velocity. A full account of nonlinear flow instability analysis based on perturbation method is very tedious and beyond the scope of this paper. Nevertheless for a given overall flow rate, the model may represent a valuable first approximation to the aggregate properties of the flow.

Nomenclature

Nomenciature

$$a = \text{sphere radius (m)}$$

$$F = \text{gravity-related parameter}$$
Fr = Froude number

$$(f, u, v, \phi, w) = \text{dimensionless variables}$$

$$G = \text{geometry-related parameter}$$

$$g = \text{acceleration (m/s^2)}$$

$$H = \text{film thickness (m)}$$

$$h = \text{dimensionless film thickness}$$
Nu = Nusselt number

$$P = \text{pressure (kg/ms^2)}$$
Pr = Prandtl number

$$p = \text{dimensionless pressure}$$

$$Q = \text{flow rate (m^3/s)}$$
Re = Reynolds number

$$(\theta, r), (x, Y), (x, y) = \text{spherical coordinates measured by angular}$$

$$displacement and radial distance$$

$$respectively$$

T = temperature (K)

Transactions of the ASME

 $T_{\rm w} =$ wall temperature (K)

(U, V) = velocity components (m/s)

$$x_{\rm s} =$$
 domain-related parameter

Greek Symbols

- α = thickness-related parameter
- $\beta =$ similarity-related parameter
- $\gamma = curvature$ -related parameter
- $\kappa =$ thermal diffusivity (m²/s)
- $\nu = \text{kinematic viscosity } (\text{m}^2/\text{s})$
- $(\xi, \eta) =$ dimensionless coordinates
 - $\rho = \text{density} (\text{kg/m}^3)$
 - $\psi =$ stream function

Subscript

0 = jet

Superscript

⁻= dimensional analysis

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